A GPU Algorithm for Convex Hull *

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Abstract

We present a novel algorithm to compute the convex hull of a point set in \mathbb{R}^3 using the graphics processing unit (GPU). By exploiting the relationship between the Voronoi diagram and the convex hull, we derive the answer from the former. Our algorithm only requires a few simple atomic operations and does not need explicit locking or any other concurrency control mechanism, thus it can maximize the parallelism available on the modern GPU.

Our implementation using the CUDA programming model on Nvidia GPUs is robust, exact, and efficient. The experiments show that it is up to an order of magnitude faster than other sequential convex hull implementations running on the CPU for inputs of millions of points. We further extend our GPU approach to obtain the Delaunay triangulation of points in \mathbb{R}^3 by computing their 4D convex hull. Our works demonstrate that the GPU can be used to solve non-trivial computational geometry problems with significant performance benefit, without sacrificing accuracy or robustness.

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1 Introduction

The convex hull $\mathcal{C}(S)$ of a set S of input points is the smallest convex polyhedron enclosing S (Figure 1). Our problem is to compute for a given set S in \mathbb{R}^3 its convex hull represented as a triangular mesh, with vertices that are points of \hat{S} , bounding the convex hull. Each point of S on the boundary of $\mathcal{C}(S)$ is called an extreme vertex. The convex hull, along with the Delaunay triangulation and the Voronoi diagram (VD) are some of the most basic yet important geometric structures. In particular, the convex hull is useful in many applications and areas of research. In scientific visualization and computer games, convex hull can be a good form of bounding volume that is useful to check for intersection or collision between objects [Liu et al. 2008; Mao and Yang 2006]. In robotics, it is used to approximate robots and obstacles for the purpose of path planning [Okada et al. 2003; Strandberg 2004]. In astronomy, it is a basic structure used to analyze the characteristics of the atmosphere [Fuentes et al. 2001; Amundson et al. 2005]. In general, convex hull is also a useful tool in biology and genetics [Wang et al. 2009] and visual pattern matching [Hahn and Han 2006].

Many CPU-based 3D convex hull algorithms have been developed and implemented over the decades. Among them, Quick-Hull [Barber et al. 1996] has been the most efficient and pop-



Figure 1: Convex hull of the Asian dragon model. Lines behind are hidden for clarity.

ular one in practice. There are also parallel convex hull algorithms [Miller and Stout 1988; Amato and Preparata 1993], but these are theoretical works with no practical implementations.

In recent years, the computational capability of the GPU has surpassed that of the CPU and is being used to solve large scale problems such as physical and biological simulation. However, this enormous power of the GPU has not been effectively exploited to solve computational geometry problems like convex hull. Existing parallel convex hull algorithms do not seem to map well to the current GPU architecture that supports tens of thousands of computational threads. The main difficulty is that such computation generally needs "global" consideration of all input data, and thus does not map well to the massively multi-threaded model of the GPU, which requires regularized work on localized data to achieve good performance. QuickHull, using an incremental insertion approach, is very difficult to be implemented efficiently on the GPU for \mathbb{R}^3 and higher dimensions, because there are many dependencies during the insertion of points.

In this report, we establish that the GPU is a useful tool to compute the convex hull in 3D with substantial speedup over sequential algorithms. The main idea of our proposed 3D GPU convex hull algorithm is to exploit the relationship between 3D Voronoi diagram and 3D convex hull so as to maximize the parallelism. From the 3D Voronoi diagram computed in digital space, we derive a good approximation which is then transformed into the desired convex hull. Our implementation uses simple data structures and does not require any explicit locking or concurrency control techniques and thus scales well with the number of cores on the GPU. We further extend our approach to compute the 3D Delaunay triangulation by obtaining it from the lower envelope of the convex hull of the points lifted to \mathbb{R}^4 .

The report is organized as follows: Section 2 reviews the related work for sequential and parallel convex hull algorithms. Section 3 outlines our 3D convex hull algorithm, while Section 4 details our implementation techniques. Some issues regarding the use of digital space computation are discussed in Section 5. Our experimental results and analysis are presented in Section 6. Section 7 extends the algorithm to compute the 3D Delaunay triangulation, and finally Section 8 concludes the report.

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2 Related Work

In this section we look at a few of the related sequential and parallel algorithms for convex hull. We also briefly discuss the star splaying algorithm [Shewchuk 2005] in \mathbb{R}^3 which we adapt to the GPU for our algorithm as explained later.

2.1 Convex Hull Algorithms for the CPU

The incremental insertion algorithm [Clarkson and Shor 1988] constructs the convex hull by inserting points incrementally using the point location technique. QuickHull [Barber et al. 1996] is a variant of such approach. Another technique is divide-and-conquer, which is used in the algorithm of Preparata and Hong [1977]. Both the incremental insertion and the divide-and-conquer approaches have a time complexity of $O(n \log n)$. In 2D and 3D, the optimal output-sensitive convex hull algorithm has a time complexity of $\Theta(n \log h)$ [Chan 1996] where h is the number of extreme vertices. There is no known efficient implementation of this algorithm. Empirically, QuickHull is found to have the same output-sensitive time complexity. Because of the good time complexity and low overhead in practice, QuickHull has been a popular approach adopted by many applications over the years.

Parallel algorithms for computational geometry in general and convex hull in particular have been extensively studied in the last few decades. For example, Miller and Stout [1988] and Amato and Preparata [1993] describe $O(\log n)$ parallel algorithms for n points using O(n) processors. These algorithms are only of theoretical interest and not yet of practical value as there are no known efficient implementations of them. One of the reasons is that these algorithms are complex, making them hard to scale on a fine-grained data-parallel massively-multithreaded architecture like the GPU. For the current multi-core systems with a small number of independent processors, algorithms designed by Dehne et al. [1995] and Gupta and Sen [2003] may be applicable, but these algorithms have no known implementations that demonstrate their uses.

2.2 Algorithms for the GPU

There has been a growing interest in using the GPU for computational geometry in recent years. Hoff et al. [1999] and Fischer and Gotsman [2006] compute the digital VD by adapting the traditional graphics pipeline of the GPU. An exact algorithm for such a computation is recently proposed by Cao et al. [2010b], using the more flexible computation model available on newer generation GPUs. Rong et al. [2008] make the first serious attempt to compute the 2D Delaunay triangulation using the GPU. Our approach of using the digital VD as an approximation to the continuous result resembles their approach. However, their work requires that the dualization produces a valid mesh, free from topological and geometrical problems, and a very subtle proof [Cao et al. 2010a] is needed for their 2D problem. Such a neat result is not known for 3D. Also, using bilateral flipping to transform an arbitrary polyhedron into its convex hull is not always possible [Joe 1989]. In another development, Jurkiewicz and Danilewski [2011] make the first successful implementation of the QuickHull algorithm on the GPU, but only for the simple 2D case.

2.3 Star Splaying in \mathbb{R}^3

Star splaying [Shewchuk 2005] is a very efficient algorithm to repair convex hulls. Here we briefly outline the algorithm in \mathbb{R}^3 which we adapt to the GPU.

In \mathbb{R}^3 , each vertex *s* of a polyhedron has a set of edges and triangles incident to it, referred to as the *star* of *s*. The set of edges opposite



Figure 2: *Star splaying algorithm. (a) A star with its link (in bold). (b) Beneath-beyond insertion.*

to s in the triangles of its star form a ring, referred to as the *link* of s (Figure 2(a)). By extending the star of s to infinity, we get a *cone*. If the polyhedron is convex, then the cone of each of its vertices is convex too, and it encloses the rest of the vertices of the polyhedron.

The stars of the vertices of a polyhedron are *consistent* with each other. That is, if the star of s contains the triangle stu, then the stars of t and u also contain this triangle. Conversely, a set of consistent stars uniquely define a surface triangulation. However, an arbitrary collection of stars not coming from a polyhedron may not be consistent with each other.

The star splaying algorithm is based on the idea that if the cones of all the points are made convex and their corresponding stars are made consistent, then these stars uniquely define the convex hull of the point set. Using the set of stars with their cones being convex, the algorithm repeatedly checks for each triangle stu in the star of s whether this triangle exists in the star of t (and u). If stu does not exist in the star of t, some points (s, u or both) will be inserted into t's star in an attempt to *splay* it to include the triangle. The insertion of a point into the star of t is done using the traditional beneathbeyond method [Edelsbrunner 1987] to guarantee that the cone of tis still convex (Figure 2(b)). If these insertions fail, implying that the triangle is enclosed by the cone of t, then some points from the link of t will be inserted into the star of s to splay it further to remove stu. In case a star of a point splays too wide that it cannot be contained in a half space, then that point is guaranteed not to be an extreme vertex. Such a star is called a *dead star*.

A nice feature of the star splaying algorithm is that creating stars having convex cones and enforcing their consistencies can both be done independently for each star. This is well suited for the parallel computation model of the GPU since stars can be checked and modified in stages without requiring any locking or concurrency control. We adapt star splaying to efficiently transform our approximation of the convex hull of S into C(S). In the following section, we describe our convex hull algorithm in more detail.

3 Algorithm Overview

The main idea of our algorithm is to utilize the relationship between the 3D Voronoi diagram and the convex hull computed from the same point set S. In particular, only the Voronoi cells of the extreme vertices of S are unbounded, i.e., extend to infinity. Thus, one can first identify these Voronoi cells to derive the extreme vertices of S. Traditionally, this observation is not computationally useful as the Voronoi diagram $\mathcal{V}(S)$ structure is harder to manage than the convex hull, and is just as expensive to compute. But, on the GPU the Parallel Banding Algorithm (PBA) [Cao et al. 2010b] can compute the digital VD very efficiently and it is a good starting point to derive an approximation of $\mathcal{C}(S)$.



Figure 3: (*a*) *Digital restricted VD.* (*b*) *Stars constructed from the digital restricted VD; they might not be consistent.*

In our algorithm, we enclose the input point set S in a large enough box \mathcal{B} which contains integer grid points, each corresponding to one unit cell of \mathcal{B} . We use the boundary (six faces) of \mathcal{B} to capture the unbounded Voronoi cells of S. Theoretically, if \mathcal{B} is large enough, dualizing $\mathcal{V}(S)$ restricted to the faces of \mathcal{B} , i.e. the *restricted VD* (Figure 3(a)), gives us $\mathcal{C}(S)$. However, since the VD we compute is in digital space, and due to the finite size of \mathcal{B} , we can only obtain an approximation of the convex hull. We apply the star splaying algorithm [Shewchuk 2005] to transform the approximation into the solution. Our algorithm can be split into five steps:

- **Step 1: Voronoi Construction.** Compute the 3D digital VD of S restricted to the boundary of \mathcal{B} . Let S' be the set of points whose Voronoi regions appear on it.
- **Step 2: Star Creation.** We first dualize the restricted VD to obtain for each point *s* in *S'* a set of neighbors, called *working set* of *s*. Then we use that to construct a *convex cone*, represented as a *star*, for the point.
- Step 3: Hull Approximation. Apply the star splaying algorithm to obtain the convex hull $\mathcal{C}(S')$.
- Step 4: Point Addition. Collect points of S that lie outside C(S') and for each of them construct its star using its nearby vertices of C(S').
- Step 5: Hull Completion. Perform star splaying again on our approximation to transform it into C(S).

3.1 Step 1: Voronoi Construction

Our aim is to approximate $\mathcal{V}(S)$ restricted to the six faces of \mathcal{B} . We first translate and then scale the input points such that their bounding box fits inside a 3D grid \mathcal{B} consisting of grid cells. Then, we compute the digital VD of S intersecting each side of the boundary of \mathcal{B} on the GPU. For each side, we project the points onto it, recording one nearest point among those that fall onto the same 2D grid cell. The two coordinates of a point are shifted to the center of the nearest 2D grid cell, while the third coordinate (the distance to the side we are projecting on) is unmodified. We then apply PBA to compute the digital VD.

3.2 Step 2: Star Creation

We dualize the restricted VD obtained in the previous step to get a set of triangles. The corners of grid cells are grid vertices, each of which is incident to a maximum of 4 different Voronoi regions. Each grid vertex incident to 3 or 4 different Voronoi regions is dualized into one or two non-intersecting triangles respectively.

Ideally, dualizing the restricted VD of S on a closed box results in a 3D polyhedron, not necessarily convex, approximating C(S').

However, in the digital restricted VD, a Voronoi cell can be, for example, disconnected, resulting in the dualized polyhedron having holes or duplicated triangles. Instead of constructing a polyhedron, we only record the information on the adjacencies of the Voronoi cells. For each triangle abc thus obtained, we add b and c to the working set of a, and similarly for b and c.

For each s in S' in parallel, we create its star (in the continuous space) from its working set such that its cone is convex (Figure 3(b)). Each GPU thread handling a point s first creates an initial star from 3 points in the working set, and then incrementally inserts the rest using the beneath-beyond algorithm.

3.3 Step 3: Hull Approximation

Star splaying is an iterative approach. Stars are repeatedly checked for inconsistency, and insertions are performed using the beneathbeyond algorithm to splay the stars when needed.

To perform the star splaying algorithm in parallel while achieving regularized work for different GPU threads, we carry out the consistency checking and the insertions of points in two separate steps, alternately performed until all the stars are consistent. For ease of implementation, the consistency checking is performed on each edge of each star in parallel, rather than on triangles. An edge is *consistent* if the stars of its two endpoints have the same two triangles incident to this edge; otherwise, it is *inconsistent*. Any inconsistency can generate up to 4 insertions. The insertion step is done by first sorting the set of insertions by the indices of the stars they are destined for, and then each parallel thread can perform the insertions into a star independent of others.

It is possible to use a CPU sequential algorithm to compute C(S'), since the set S' derived from the digital restricted VD is quite small. However, since the stars constructed from the previous step are already very close to C(S'), a parallel implementation of the star splaying algorithm on the GPU gives much better performance.

3.4 Step 4: Point Addition

Due to $\mathcal{C}(S')$ being an approximation, there might be extreme vertices of S that are missing in it. We use $\mathcal{C}(S')$ to check the points in S and remove those that are inside the hull. The rest of the points can potentially be extreme vertices. This is the reason why we perform star splaying in Step 3. We first perform the checking in digital space by rendering the triangles of $\mathcal{C}(S')$ with the view direction orthogonal to each side of \mathcal{B} in turn. Then, we use a depth test to eliminate points that clearly lie inside $\mathcal{C}(S')$. Each GPU thread handling a point s in S projects s onto each side of \mathcal{B} and compares its depth value d_s with the value d on the depth map on that side (with depth value increasing in the viewing direction). If $d_s - d \leq \tau$ where τ denotes the threshold, then s is potentially an extreme vertex. The depth test is done in digital space, so a conservative threshold (which is equal to 1 pixel width) is used to safely remove non-extreme vertices; see Section 5.1.

To further eliminate non-extreme vertices, we perform another round of checking in continuous space. For each point s that passes the depth test, we also record a triangle A that covers its projection in one of the viewing directions. Notice that A is close to s. Pick an arbitrary point r on C(S'). The point s is either beyond one of the triangles in the star of r, or the ray \vec{rs} intersects C(S') at a triangle B. Using a technique similar to point location, starting from A we can quickly find B, and accurately determine whether s is inside or outside C(S'). If s is outside, we use B to form a star for s, otherwise it is eliminated. All this computation can be done on each point independently in parallel. The new stars together with C(S')form an approximation of C(S).

3.5 Step 5: Hull Completion

Our approximation now contains all possible extreme vertices of S and a few more. By performing star splaying again, as detailed in Step 3, we transform our approximation into the convex hull of S.

4 Implementation Details

We implement our algorithm using the CUDA programming model and OpenGL on Nvidia GPUs. Due to the nature of the GPU, it is more efficient to allocate memory in large chunks rather than dynamically allocate many small blocks. For our implementation, we use two lists to store the description of the stars and their edges, called the *star list* and the *edge list*, as shown in Figure 4. Each star has a contiguous chunk of memory whose size is enough to store all of its current edges plus a certain amount of free space. Each star records the coordinates of its point, its status (whether it is dead or not), the number of edges, the size of memory allocated for it, and the starting location of its storage in the global edge list. Each edge of a star records the index of the other endpoint, and a flag used when checking for consistency. The edge list of a star represents its link vertices in counter-clockwise order.

The difficulty here is that the edge list has a dynamic size as stars are shrinking as well as expanding during the star splaying process. Any time a star uses up its chunk of allocated storage, we have to expand the edge list. When such an expansion occurs we also use the opportunity to shrink or expand the storage of all the stars to maintain a "healthy" ratio (say 1.2) of available storage for each. This helps to reduce the number of times we need to reallocate the edge list. Also, since we start star splaying on a good approximation of the convex hull, the stars typically do not grow drastically.



Figure 4: Data structures for stars and edges. \times indicates a dead star.

4.1 Step 1: Voronoi Construction

Before applying the PBA, we need to project the points on the six sides of the box \mathcal{B} . This operation entails a lot of random atomic memory access to the global memory that is highly inefficient on the GPU. Instead, we perform all the projections on the GPU shared memory to speed up this step.

 \mathcal{B} is partitioned into bricks, each of size $k \times k \times k$. For each point in S, we find the brick that encloses it using a CUDA kernel. We accumulate the points that belong to the same brick into a contiguous chunk using CUDA radix sort. We identify the starting offset of each such chunk in the sorted list using another CUDA kernel.

We use six textures to store the projections on the six sides of \mathcal{B} . For each $k \times k$ tile of a texture, we use a block of CUDA threads to process the points enclosed in the bricks that project onto this tile. When many points project onto a single pixel, we store the point closest to the tile by using the *atomic minimum* operation. This is applied on a shared memory array of the block and thus is highly efficient compared to using it on global memory. k can typically be chosen to be 32, so that the $k \times k$ tile can fit in shared memory. Though we use floating point for the point-to-tile distance, we can still use the CUDA integer atomic minimum operation. This is because positive floating point numbers can be compared as integers, with the same binary representation, without affecting their order. The result of these projections is coherently written out to global memory to apply PBA and obtain the restricted VD.

4.2 Step 2: Star Creation

We construct the working set for each point by scanning the resulting VD textures constructed in the previous step. For each triangle identified, we generate 6 pairs, each pair (a, b) indicating that b is in the working set of a. First, we let one CUDA kernel count the number of pairs generated by each grid corner. Next, we pre-allocate an array to store the pairs, and use CUDA parallel prefix sum primitive to compute for each corner the offset in the array to store its pairs. After that, we call another kernel to scan the textures again, generating the working set pairs. Lastly, we sort the list of pairs using CUDA radix sort, remove duplicates and identify the working set for each point as a contiguous chunk of pairs.

Based on the working sets thus constructed, we allocate the storage for the star list and the edge list. A CUDA kernel is used to construct an initial star consisting of 3 link points for every point in S'. Each CUDA thread constructing an initial star takes 3 points from its working set, checks the 3D orientation and stores these points in the edge list of that star in counter-clockwise order.

After that, the rest of the working set of each point is inserted into its star in a single kernel. Each CUDA thread processes the working set of a point, independent of other points. For each insertion of t into s, we go through the star of s, identifying a (continuous) series of beneath triangles, removing their corresponding edges and inserting t into the edge list of s accordingly.

The beneath-beyond insertion relies heavily on the 3D orientation predicate. It is important that the predicate is computed exactly and co-planar cases are handled correctly. More importantly, the predicate should give the same result when checked from different stars for the star splaying algorithm to converge. In order to achieve this, all our predicates are performed with the Simulation of Simplicity (SoS) technique [Edelsbrunner and Mücke 1990] and exact arithmetic [Shewchuk 1997]. The flexibility of the CUDA programming model makes such complex computation possible on the GPU.

4.3 Step 3: Hull Approximation

Algorithm 1 Star splaying on the GPU

- 1: flag all edges to be checked for consistency
- 2: repeat
- 3: collect the edges that are flagged
- 4: allocate space for possible insertions
- 5: check the flagged edges and generate insertions
- 6: sort and compact the list of insertions
- 7: **if** a star needs more space **then**
- 8: expand the edge list
- 9: perform the insertions to splay stars
- 10: flag edges that need to be checked in the next iteration
- 11: until there are no more flagged edges

The pseudocode of the star splaying implementation on GPU is outlined in Algorithm 1. In Line 3, we do a parallel stream compaction on the edge flags to obtain the list of edges to be checked for consistency. Each inconsistent edge can potentially lead to up to four insertions into different stars (see Section 2.3). We pre-allocate storage for these possible insertions in Line 4. In Line 5, we use a CUDA kernel where each thread processes one edge. The insertions are sorted and compacted in Line 6 and duplicates are removed. Each star then checks if it has enough free space in its edge list and the edge list is expanded if needed (Line 7 and Line 8). This expansion is done by computing the required space for each star using a kernel, allocating a new edge list, and then copying the edges over. The insertions (Line 9) are performed similar to those in Step 2. In Line 10, we flag all newly created edges. Also, during the insertions, if an edge ab in the star of a is deleted, then the edge ba in the star of b, if any, needs to be flagged too.

4.4 Step 4: Point Addition

The first round of checking in this step is carried out in OpenGL, which works seamlessly with other steps done in CUDA. As we keep edges rather than triangles, we first use a CUDA kernel to generate a list of triangles in C(S') from the stars. To avoid generating duplicate triangles, each triangle *abc* is created only by the star of *a* where *a* has the smallest index among the three. Similar to other steps, we first count, then use parallel prefix sum to compute the offset before actually generating the triangle list.

When a triangle is rendered, we record in the color buffer the index of one of the three vertices so that we can use it as the starting point for our point location in the second round of checking. After the rendering, the depth buffer is processed by a CUDA kernel. Each thread processing a point in S - S' checks the depth value to see whether the point can potentially be outside or not. If outside, this point becomes a candidate for the next round of checking.

In the second round of checking, we use one CUDA thread to check one candidate found in the previous round. Let the candidate be sand the corresponding point recorded at the projection of s in the color buffer be c. Also, let r be an arbitrary point in S' where $r \neq c$. In order to determine the triangle B in $\mathcal{C}(S')$ that is intersected by the ray $r\bar{s}$, we start walking from c. Each vertex t on the link of c together with the line rc form a plane, and we are interested in the half-plane defined by rc that contains t. The collection of these half-planes partitions the space into several unbounded subspaces around rc; one of such subspaces contains s, which can be identified using 3D orientation checks. This subspace tells us which vertex on the link of c gets us closer to s, until we reach B. After that, using one more 3D orientation check, we can determine accurately if s is outside $\mathcal{C}(S')$, in which case the three vertices of Bform the initial star of s.

5 Digital Approximation Issues

In this section, we discuss some issues of the use of digital space computation to approximate that in the continuous space.

5.1 Digital Depth Test

In Step 4, we use the six sides of the boundary of \mathcal{B} as the viewing planes. We compare the depth d_s of each point s with the minimum depth value of $\mathcal{C}(S')$ at the corresponding projection of s to quickly exclude points that are inside $\mathcal{C}(S')$. However, since the depth buffer we obtain when rendering $\mathcal{C}(S')$ is of finite resolution, the depth value d of the projection of s is actually the depth value of the center of the cell containing this projection. Depending on the triangle covering that projection, $(d_s - d)$ can be arbitrarily large; see Figure 5. The following claim shows that as long as we keep every point s that has $(d_s - d) < 1$ in one of the projections, we do not miss any point outside $\mathcal{C}(S')$. This tight bound allows us to throw away most of the points that are inside $\mathcal{C}(S')$.

Claim 1. Let $s \in S - S'$ be a point outside $\mathcal{C}(S')$. In (at least) one of the six renderings of $\mathcal{C}(S')$ orthogonal to a side of \mathcal{B} , we have



Figure 5: *The digital depth test of a point s against a triangle T on the boundary of* C(S') *when s is outside* C(S')*.*

 $(d_s - d) \leq \tau$ where $\tau = 1$ pixel width.

Proof. The point s is inside a unit cell of \mathcal{B} whose center is the grid point $(\bar{x}, \bar{y}, \bar{z})$. The coordinate of s is $(\bar{x} + \delta_x, \bar{y} + \delta_y, \bar{z} + \delta_z)$ where $\delta_x, \delta_y, \delta_z \in [-0.5, 0.5]$. Let T be the triangle covering the cells containing the projections of s in different viewing directions, and the plane equation of T be ax + by + cz + K = 0. Without loss of generality we assume that $a \ge b \ge c$.

Since T appears in the depth buffer, and C(S') is convex, T must be visible from three different viewing directions. This forms a coordinate system in which the plane equation of T has $a, b, c \ge 0$. In the viewing direction along the positive x-axis, $d_s = \bar{x} + \delta_x$ and d is the depth of T at (\bar{y}, \bar{z}) . As s is outside C(S') and thus is in front of the plane of T, $a(\bar{x}+\delta_x)+b(\bar{y}+\delta_y)+c(\bar{z}+\delta_z)+K \le 0$, and we thus have:

$$\begin{aligned} d_s - d &= (\bar{x} + \delta_x) - \left(-\frac{b\bar{y} + c\bar{z} + K}{a} \right) \\ &= \frac{a(\bar{x} + \delta_x) + b(\bar{y} + \delta_y) + c(\bar{z} + \delta_z) + K}{a} - \frac{b\delta_y}{a} - \frac{c\delta_z}{a} \\ &\leq \frac{b}{2a} + \frac{c}{2a} \leq 1 \end{aligned}$$

It is possible that the depth values of s used in checking in the six viewing directions belong to different triangles. Suppose that the depth value of triangle T is used in one of the directions, then from the above argument, there is one direction in which the depth d of the plane containing T fulfills the inequality $(d_s - d) \leq 1$. Suppose T' is the other triangle that covers s in that direction, then due to the convexity of C(S'), the depth d' of T' must be no smaller than d, and thus $(d_s - d') \leq 1$, as required.

5.2 Convex Hull Approximation

Due to the nature of digital space and that of our approach, there are three issues that can affect the performance of our algorithm: *slicing problem*, *under-approximation problem* and *overapproximation problem*; see Figure 6.

Slicing problem. This problem is the result of using a bounded box \overline{B} to find the Voronoi cells that are unbounded. As some of the bounded cells can extend beyond \overline{B} , they are captured although they do not correspond to extreme vertices. Figure 6(a) shows a 2D example where among the five cells being captured, only those of the round white points are unbounded. To reduce the number of wrongly captured Voronoi cells, we scale the point set to a slightly smaller volume inside \overline{B} when performing Step 1.

Under-approximation problem. When we have multiple points projected to the same pixel, we can only record one point, and thus there are potentially many more points outside C(S'). See Figure 6(b) for a 2D illustration where the round black points are kept, the solid line denotes part of C(S') and the dashed line denotes part of C(S). The round white points are missing points, many of which



Figure 6: Three problems associated with the computation in digital space.

are outside $\mathcal{C}(S')$. By using a very efficient depth test in Step 4 of our implementation and accurate location of a nearby triangle for every point outside $\mathcal{C}(S')$, we are able to construct a very good star for that point. This reduces the amount of splaying needed in Step 5.

Over-approximation problem. This problem is caused by the shifting of points in Step 1. In certain cases, for example when points are distributed near the surface of a cube axis-aligned with \mathcal{B} , many points that are not extreme vertices are shifted outside and are legitimately captured in Step 1. See Figure 6(c) for a 2D illustration, where after Step 1 all the round black points, after shifted to the square black grid points, are captured. In our implementation, for each side of \mathcal{B} we only shift 2 coordinates of the points while keeping the third one untouched. This produces a much better approximation of the restricted VD and thus reduces the effect of this problem.

6 Experimental Results

Our algorithm is implemented using the CUDA programming model by Nvidia, and can easily be ported to OpenCL. All the experiments are conducted on a PC with an Intel i7 2600K 3.4GHz CPU, 16GB of DDR3 RAM and an Nvidia GTX 580 Fermi graphics card with 3GB of video memory, unless otherwise stated. Visual Studio 2008 and CUDA 4.0 Toolkit are used to compile all the programs, with all optimizations enabled. We compare the performance of our implementation, called gHull, with the two fastest sequential implementations of the Quickhull algorithm: Qhull and CGAL. Qhull handles roundoff errors from floating point arithmetic by generating a convex hull with "thick" facets: any exact convex hull must lie between the inner and outer plane of the output facets. On the other hand, CGAL uses exact arithmetic, which is similar to our implementation. In our experiment, we found that CGAL always runs slower than Qhull due to its use of exact arithmetic.

All the results below of gHull are based on the same set of parameters: grid size 1024^2 , while point set is scaled to 80% of the volume of \mathcal{B} . The rendering buffer in Step 4 is of size 512^2 . Using bigger grid size gives better approximation at the cost of slower VD computation, so it gives little running time improvement. A larger buffer for the depth test is also not needed, since it incurs extra rendering cost.

Representative Data. We generate points randomly with coordinates between [0.0, 1.0]. Points are distributed uniformly in four distributions: a cube, a ball of radius 0.5, a box with thickness of 0.01, and a sphere with thickness of 0.01. The cube distribution has very few points on the convex hull, while many points inside can easily be removed by the Quickhull algorithm. The ball distribution is similar, but with a bit more points on the convex hull. The box distribution also has very few extreme vertices, but points are distributed close to the convex hull, so it is harder to eliminate them. The sphere is the extreme case where many points are on the convex hull.



Figure 7: Speedup of gHull over Qhull and CGAL.

vex hull, while the rest of them are also close to it. These synthetic test cases are highly representative for the convex hull problem.

The speedup of gHull are presented in Figure 7. In general, gHull is 2x to 10x faster than Qhull, and is 2x to 40x faster than CGAL for the first three distributions. Notably, for the sphere distribution where not only are there many extreme vertices, there are also many points close to the convex hull, gHull is up to 90x faster than CGAL and up to 30x faster than Qhull, even with all the computation being exact. This is mainly because our digital restricted VD gives a very good approximation, which our star splaying implementation can quickly transform to a convex hull, and Step 4 of our algorithm is also very fast in eliminating non-extreme vertices. As an example, for 10^7 points in the sphere distribution, Step 2 returns around 54,000 points, of which Step 3 keeps 32,000. Step 4 then adds 57,000 points, and the final convex hull after Step 5 has 55,000 points. These are very small numbers compared to the number of input points. Similar performance can be observed for other point distributions, as Figure 8 shows similar running time of gHull for the same number of points.

Sensitivity Analysis. The use of the digital approximation is affected by how close the points are to each other and to the convex hull. We investigate this in Figure 9(a). Here we show the running time and speedup of gHull on the sphere distribution with thickness varying from 0.5 (a ball) to 0.0001. The running time increases as the sphere gets thinner, especially at thickness 0.0001 which is only 0.1 pixel width given that we use a 1024^2 grid size. The speedup over the CPU implementations initially increases as the Quickhull algorithm becomes less effective in eliminating non-extreme vertices, then decreases but is still 10x faster.

Process Analysis. Figure 9(b) shows the running time of each step of our algorithm for 10^7 points. As expected, the behavior differs on different distributions. While the running time of Step 1 and Step 4 remains the same since it is not affected by how the points are distributed, the running time of other steps varies significantly.





Figure 8: Running time of gHull on different test cases.



Figure 9: Running time of gHull on (a) spheres of different thickness, (b) different steps and (c) GPUs with different number of cores.

Step 3 takes more time on the box distribution due to the overapproximation problem, while Step 5 takes more time on the sphere distribution due to the under-approximation problem. Step 2 only takes a small portion of running time for all distributions.

Scalability. Since our algorithm is designed for regularized computation on localized data, it is expected that it scales well with the number of cores on the GPU. This is confirmed in Figure 9(c). Here we run gHull on 10^7 points on different graphics cards: a GTS 450 with 192 cores, a GTX 460 with 336 cores and a GTX 580 with 512 cores. Clearly our implementation scales well with the number of cores, achieving close to 3x speedup when the number of cores increases by 3x, in all the different distributions. When we normalize the running time on different cards by multiplying it with the number of cores and their clock-speed (1.7 GHz, 1.4 GHz and 1.6 GHz respectively), we get roughly the same results, with newer generation GPUs being slightly faster than the older ones.

Popular 3D Models. See Figure 1 and Table 1. We use models of over a million points from the Stanford 3D scanning repository. Points of these models are densely distributed on the surface, while their convex hulls have very few points (only around 1,000). Nevertheless, gHull still manages to out-perform Qhull 2x to 6x and CGAL by 5x to 15x.

Model	# Points	Running time (ms)		
	(millions)	Qhull	CGAL	gHull
Asian dragon	3.6	540	1181	150
Thai statue	5.0	692	1538	168
Lucy	13.9	1884	4488	266

Table 1: Running time on different 3D models.

Limitation. While allowing us to perform most processing in parallel with regularized work and localized data, one limitation of using the digital space to approximate the computation in the continuous space is its uniformity. It is possible to design a test case where points are badly distributed (e.g. points arranged on a thin line convoluted in the space), resulting in a bad digital approximation, and thus lower overall performance. Such a case, however, is not common in practice.

7 Delaunay Triangulation in \mathbb{R}^3

Given a set S of input points in \mathbb{R}^3 , its *Delaunay triangulation* (DT) $\mathcal{T}(S)$ is a triangulation of S such that the circumsphere of each tetrahedron contains no other point of S. Figure 10(a) is an example of the 3D DT. The DT in \mathbb{R}^3 can be computed by lifting the points to \mathbb{R}^4 on a parabolic map [Edelsbrunner and Seidel 1985], and then computing the lower envelope of their convex hull.

Using an approach similar to our convex hull algorithm, the 3D DT can be constructed using the GPU in three steps: compute a 3D

digital VD and dualize it to get a digital approximation of the DT; augment the approximation with the rest of the points that we miss in the digital space; and finally use star splaying to transform the approximation into the final solution.

The dualization of the 3D digital VD is not as straightforward as in the 3D convex hull, since a 3D grid vertex can be incident to up to 8 Voronoi regions. Instead we use the *perturbed grid* interpretation of Bendich et al. [2010] to perform the dualization. With such a perturbation (Figure 10(b) and 10(c)), only four regions can meet at a perturbed grid vertex. The perturbation is only a guide for the dualization process, needing no extra computation.

Unlike the 3D convex hull problem, here every missing point in the digital approximation needs to be added back. For every point s' that appears in the digital VD, we create a missing set $\mathcal{M}_{s'}$ consisting of the missing points that are mapped to the same grid point as s'. Then, the working set of a missing point in $\mathcal{M}_{s'}$ is set to be the working set of s' plus all other points in $\mathcal{M}_{s'}$.

A major complication of the 4D convex hull computation is the exceedingly complex computation involved in the exact 4D orientation predicate with SoS. To regularize the work of CUDA threads in kernels that involve such computation, each of these kernels are separated into two, one to perform a fast check, and only those checks that cannot be decided need to go through the second kernel where the exact SoS predicates are performed. This means that the insertions into a star can no longer be done by one thread in a single kernel. Instead the insertions are performed in several rounds, in each of which only one point can be inserted into a star.

Our implementation of the 3D DT computation on the GPU is very robust, and can handle test data of millions of points. When the points are uniformly randomly distributed, our program outperforms CGAL, the fastest 3D DT implementation according to our experiments, by 2x to 3x. We are working on improving our implementation to achieve further speedup.



Figure 10: (*a*) A cut away 3D DT. (*b*) A grid of 8 perturbed cells. (*c*) A perturbed cell marked with its 24 grid vertices.

8 Concluding Remarks

This report introduces an algorithm to compute the convex hull in 3D using the GPU. By first computing an approximation of the solution, our algorithm converts the given problem into one that is easier to process concurrently using a massive number of threads performing regularized work on localized data. With a careful design, our implementation is efficient yet remains exact, robust and scalable to the number of GPU cores. Our experiment on different test cases shows that our implementation on CUDA is an order of magnitude faster than the best sequential CPU implementations. This demonstrates yet another non-trivial computational geometry problem that can be solved efficiently using the GPU.

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